

Binary Neutron Stars & The Peta-scale Challenge

David Nielsen
Brigham Young University

AEI Potsdam-Golm
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Collaborators

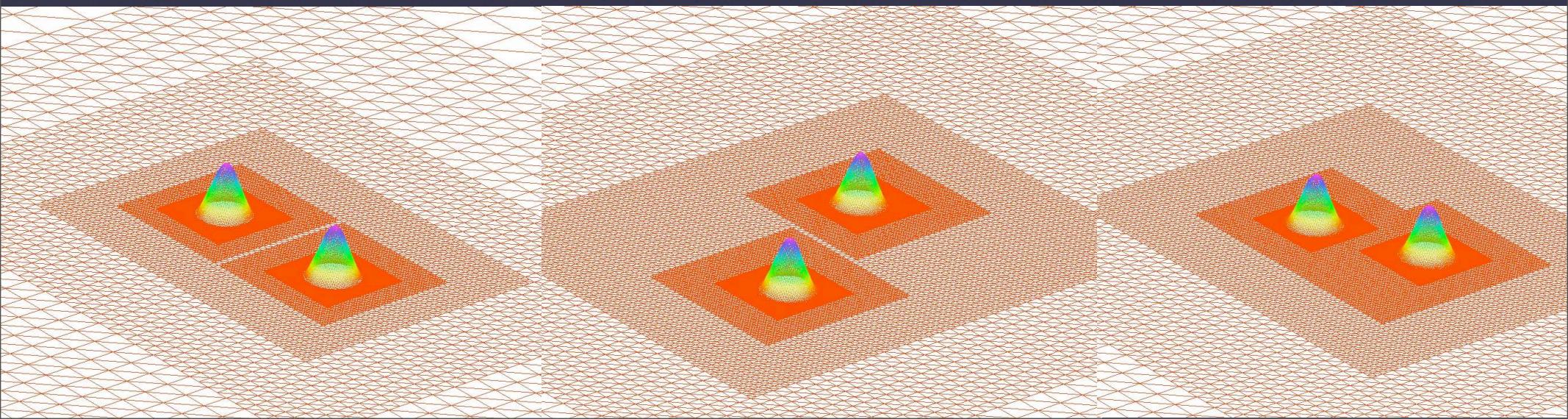
- Matthew Anderson (LSU > BYU)
- Eric Hirschmann (BYU)
- Luis Lehner (LSU)
- Steve Liebling (LIU)
- Patrick Motl (LSU)
- Carlos Palenzuela (LSU > AEI)
- Joel Tohline (LSU)

Outline

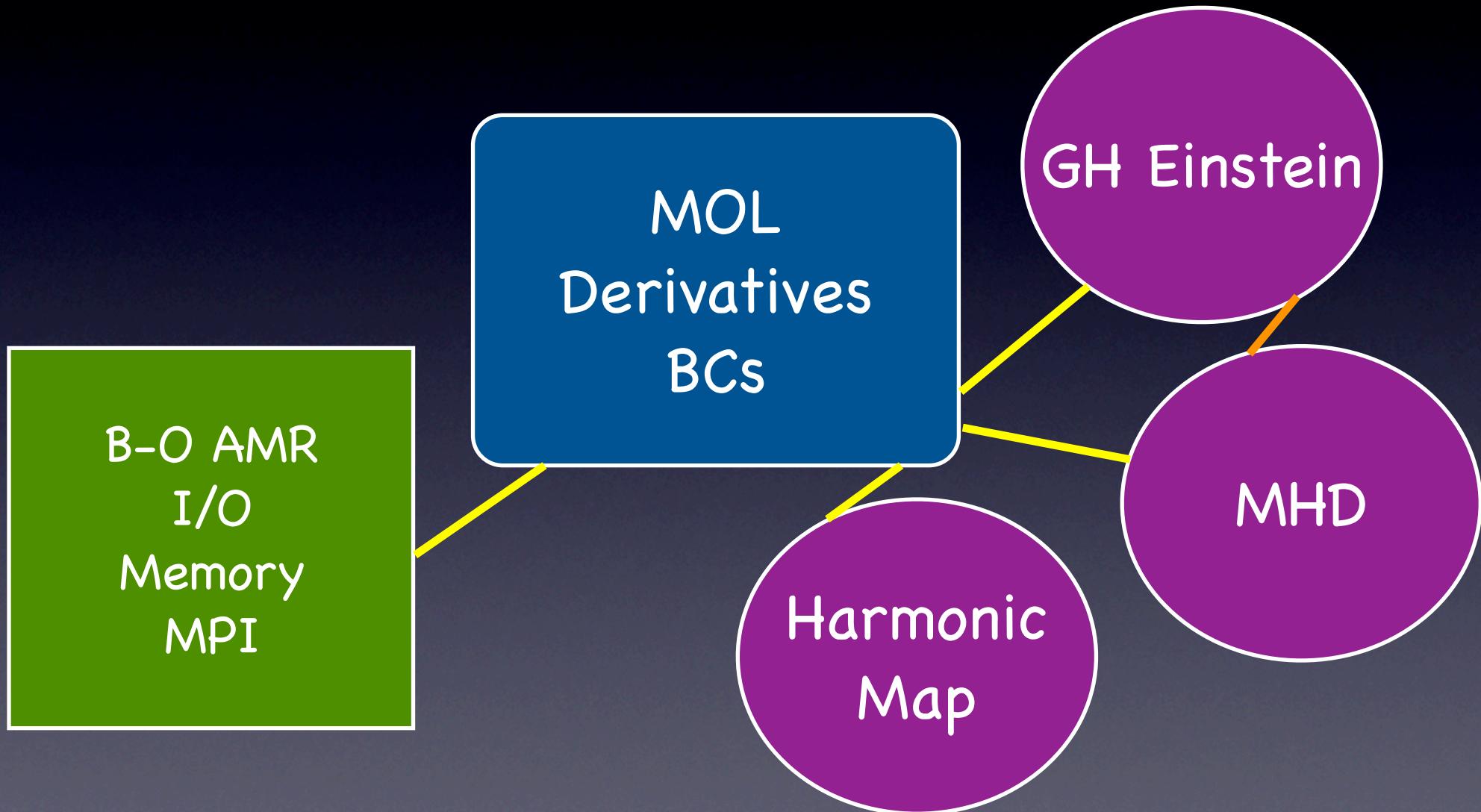
- Introduction to our GRMHD code
 - HAD
 - GRMHD
- Initial tests
- Binary neutron stars
- Binary mergers & magnetic fields

Methods: HAD

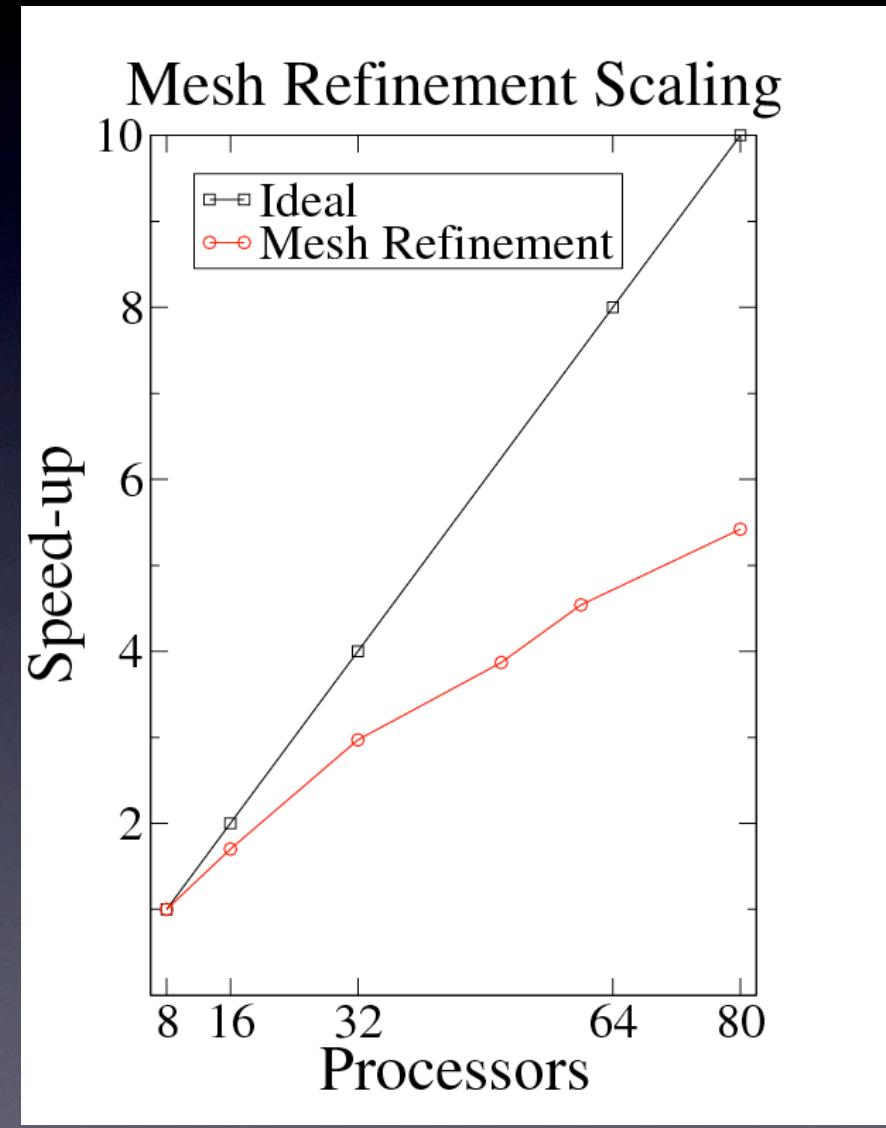
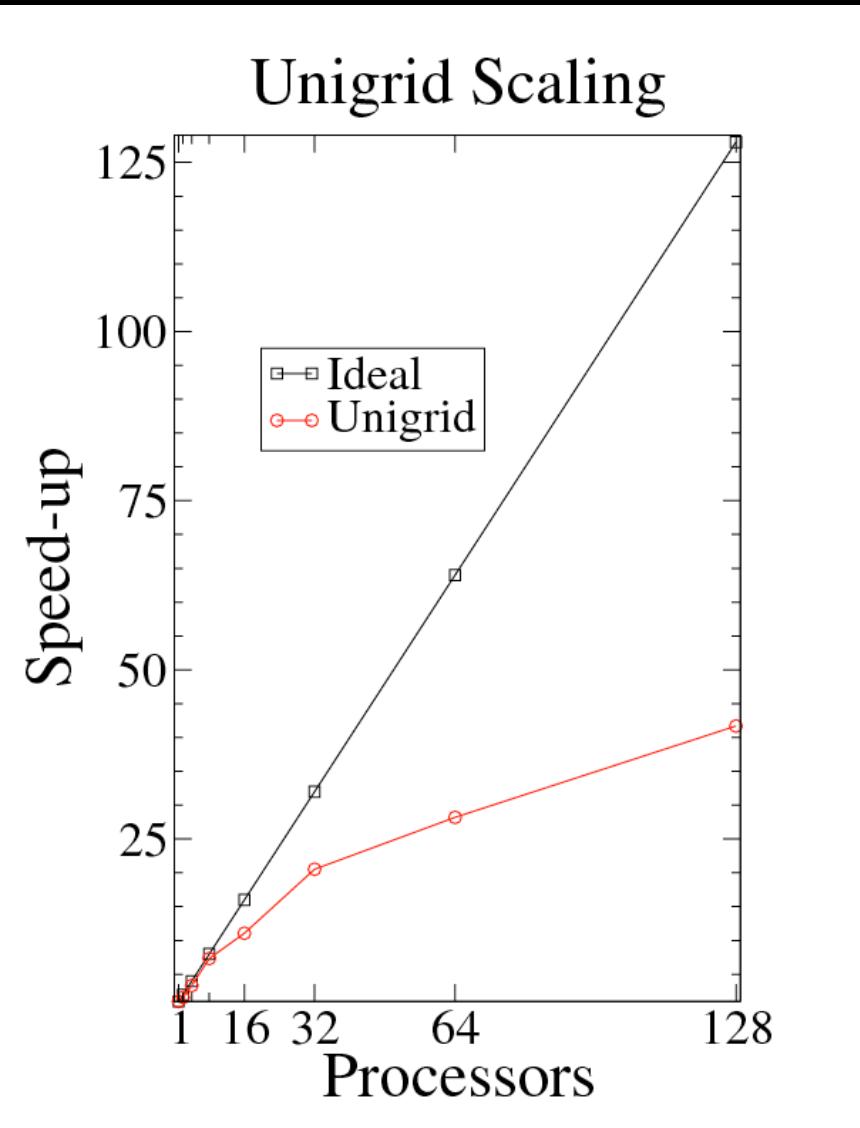
- Distributed parallelism with MPI
- Berger-Oliger AMR
- Hyperbolic & Elliptic equations
- Finite Difference & Finite Volume
- Tapered Boundary Conditions
- Arbitrary accuracy (3rd, 4th order)
- Refinement using Shadow Hierarchy or user-defined criteria
- F77 (with an assist from Perl) ... fast



HAD (cartoon form)



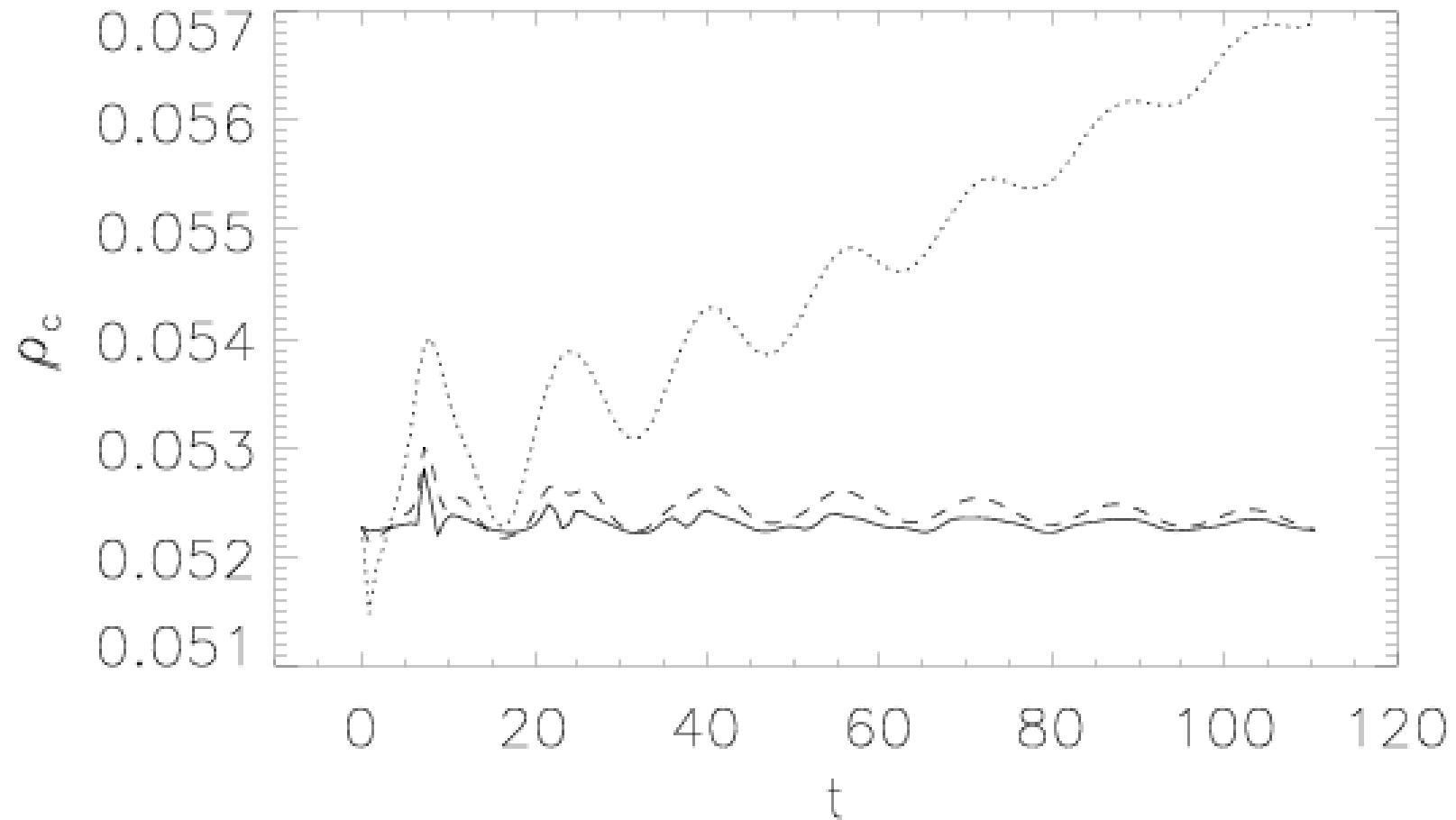
HAD: Strong Scaling Tests



Methods: GRMHD

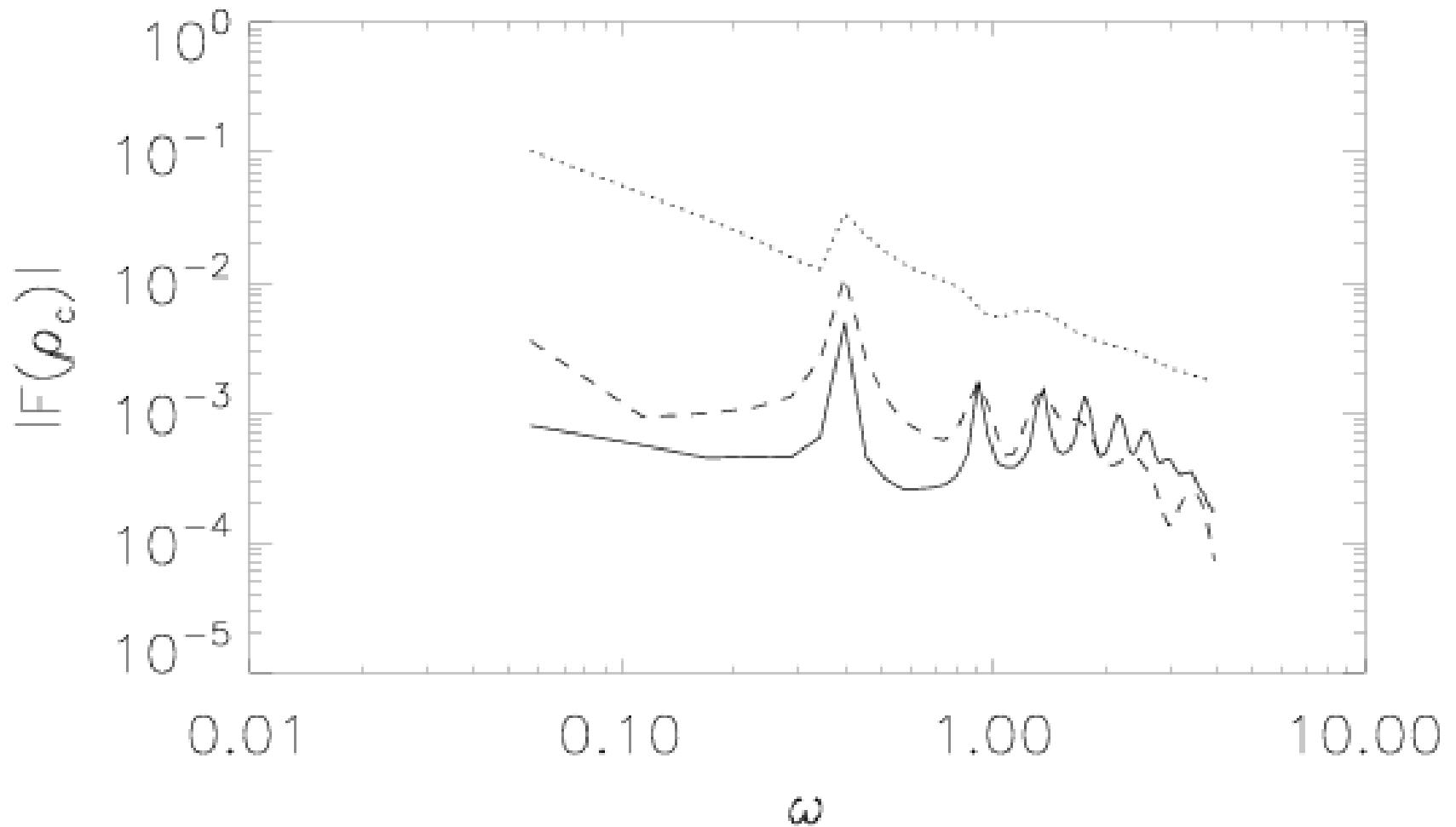
- MHD
 - CENO (FD, central-upwind) & Method of lines
 - HLLE numerical flux
 - PPM reconstruction
 - WENO interpolation for AMR
 - Hyperbolic divergence cleaning
 - Polytropic equation of state
- GR
 - Generalized Harmonic formulation
 - First order form
 - Constraint preserving boundary conditions
 - Difference operators satisfy Summation By Parts
 - Kreiss-Oliger dissipation

Single Stars: Central Density



TOV star, domain [-10,10], R=1.01, M=0.09, Gamma=2, K=1
Resolutions: 20, 40 & 80 points/star

Single stars: Fourier modes

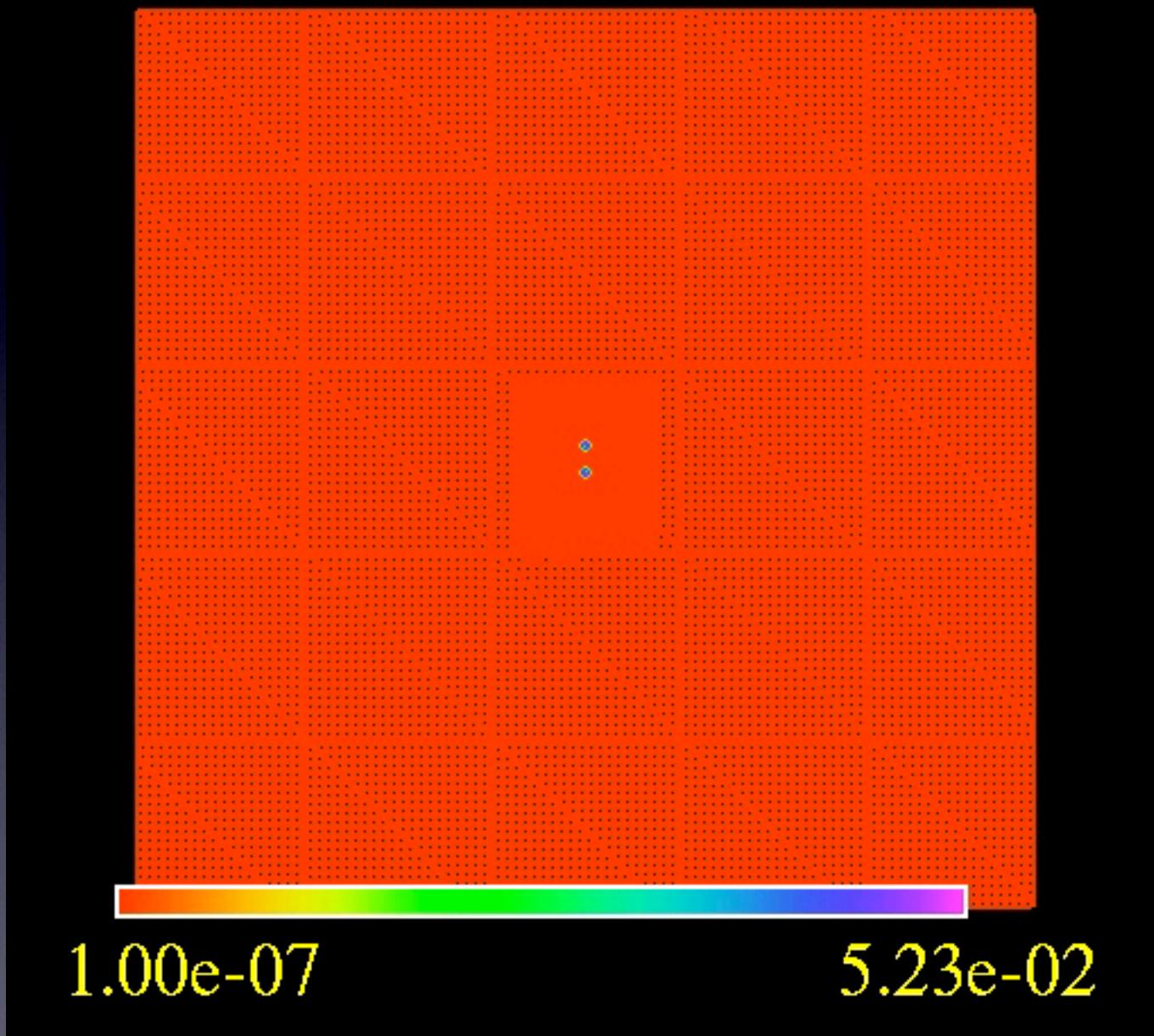


Fundamental: 14,400 Hz (Gamma=2, K=1)

Orbits

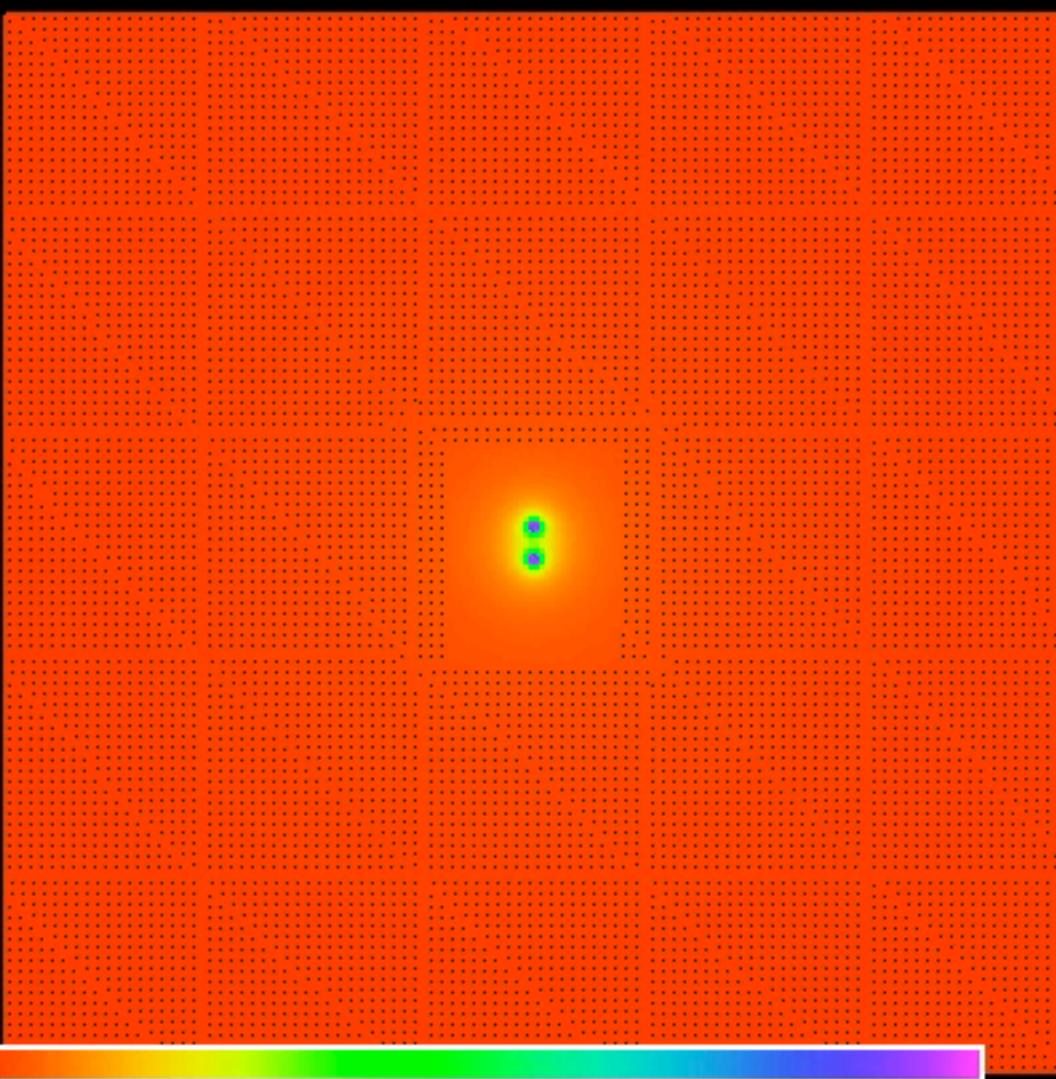
- Initial data using Kerr-Schild superposition
- $M=0.09$, $R=1$
- Domain $[-100, 100]$, $d = 6$, $v=0.09$
- Resolution: 32 pts/star
- GW extraction at 50, 60, 70

t=0.00



$t=0.00$

g_{xx}

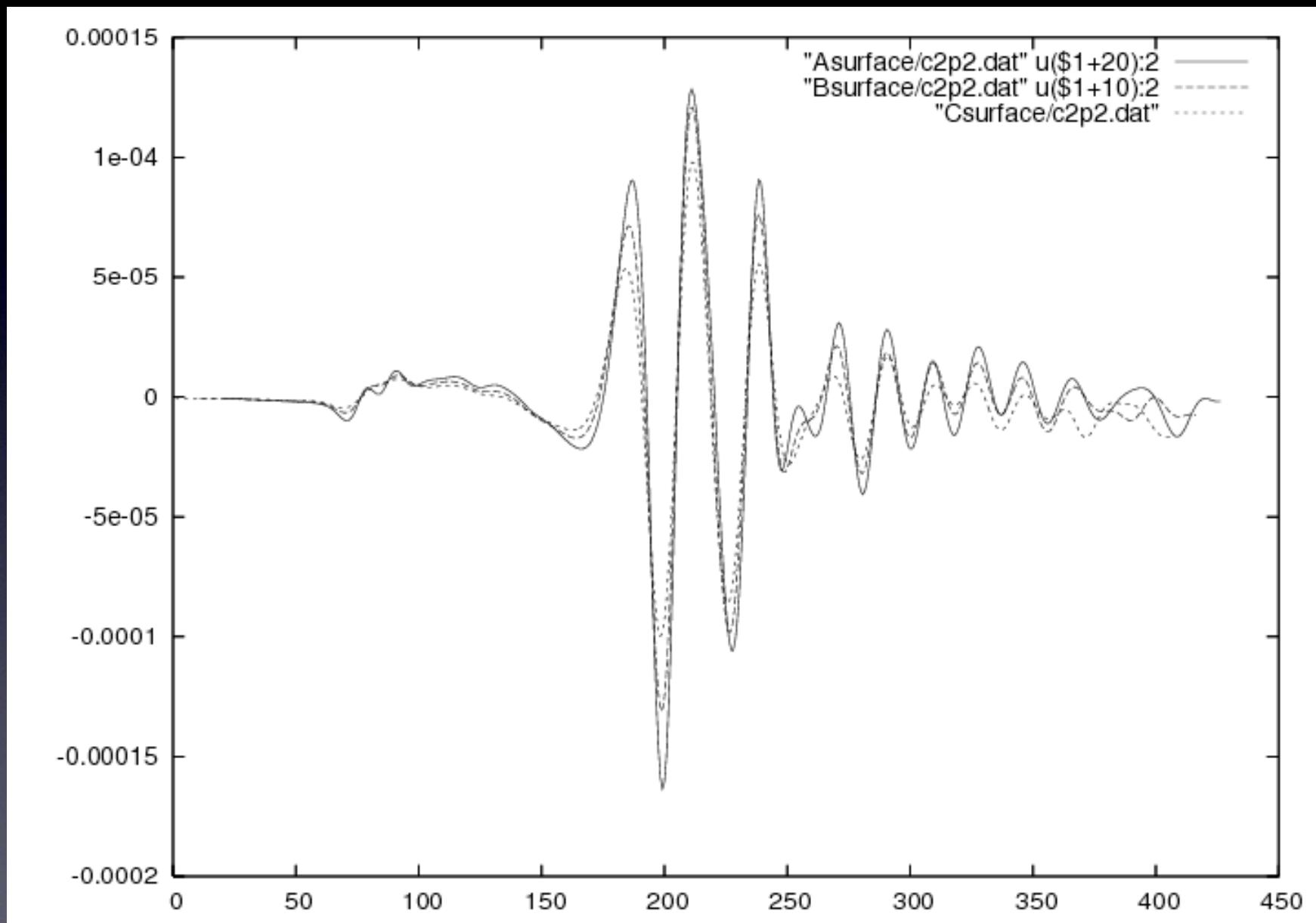


$1.00\text{e}+00$

$1.46\text{e}+00$

$[-100.00, 100.00], [-100.00, 100.00]$

Wave Extraction



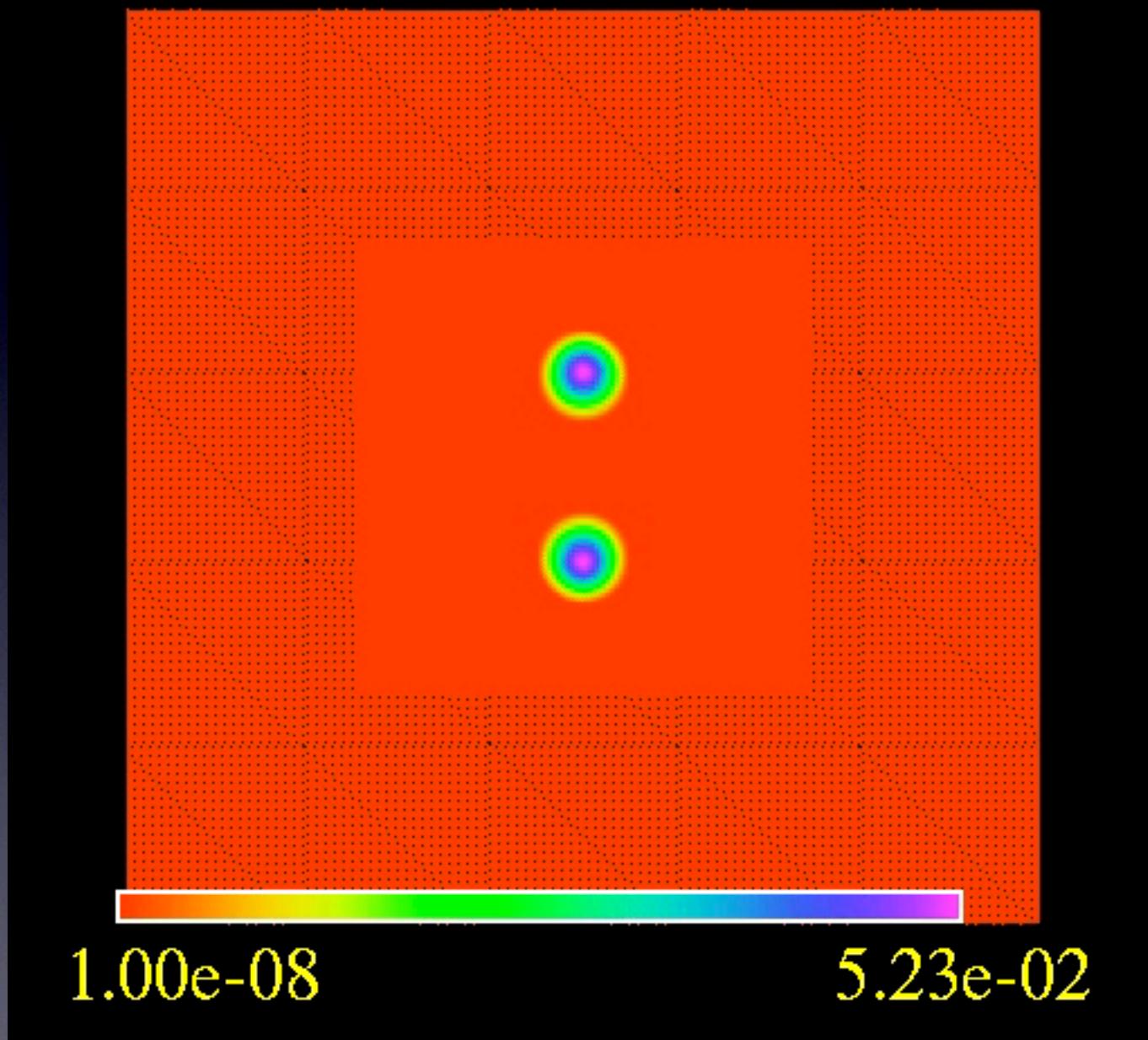
The $l=2, m=2$ mode. Extraction at $R = 50, 60 \& 70 R$.

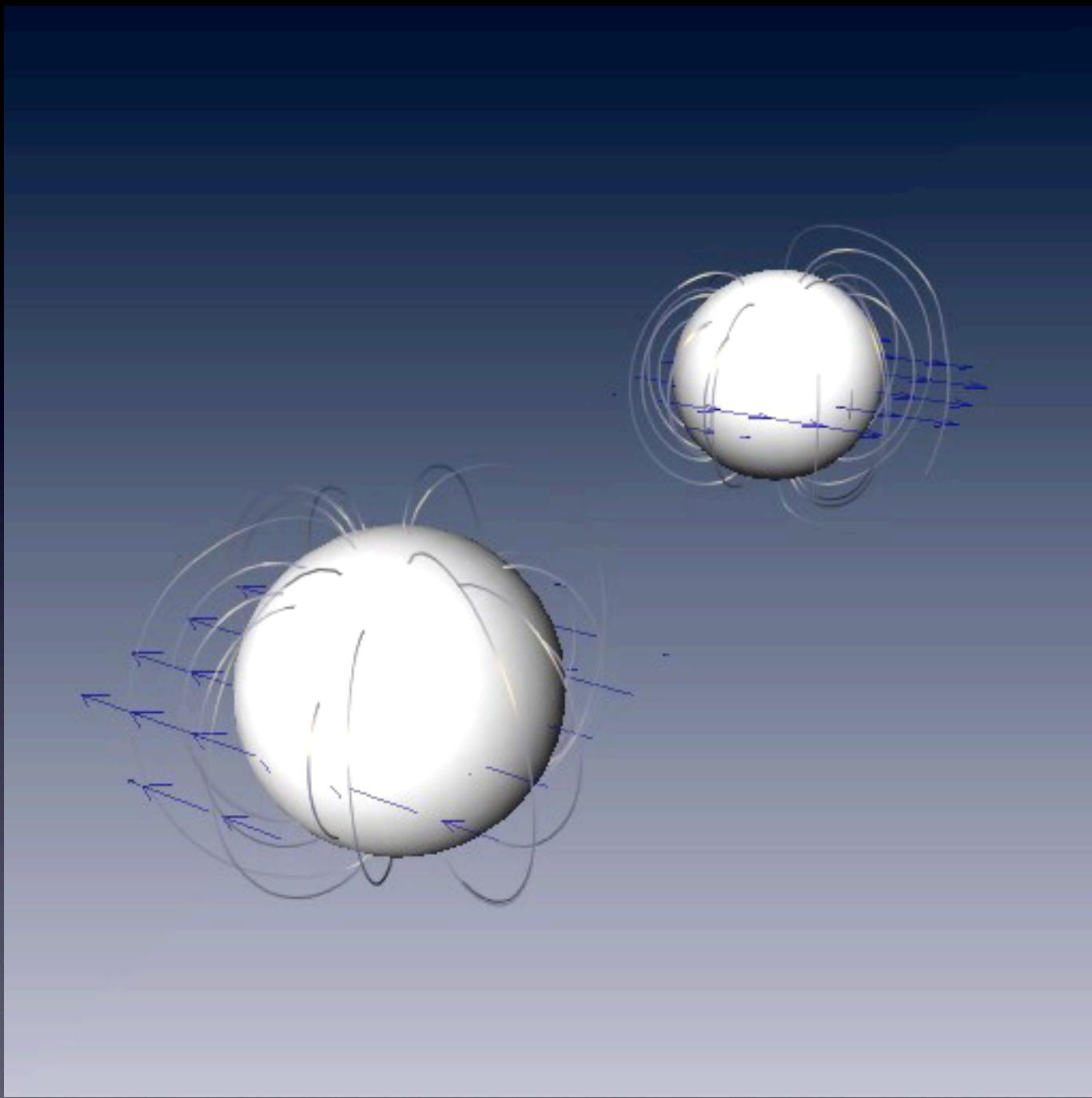
Mergers with Magnetic Fields

Sotani, Yoshida, Kokkotas, PRD 75 084015 (2007)

- Initially no differential rotation
- Domain [-10,10]
- $R = 1.01$, $M=0.1$
- Initial separation, $d=4$
- Initial boost, $v=0.08$
- 20 pts/star
- $\Gamma=2$, $K=1$
- 2 levels of refinement

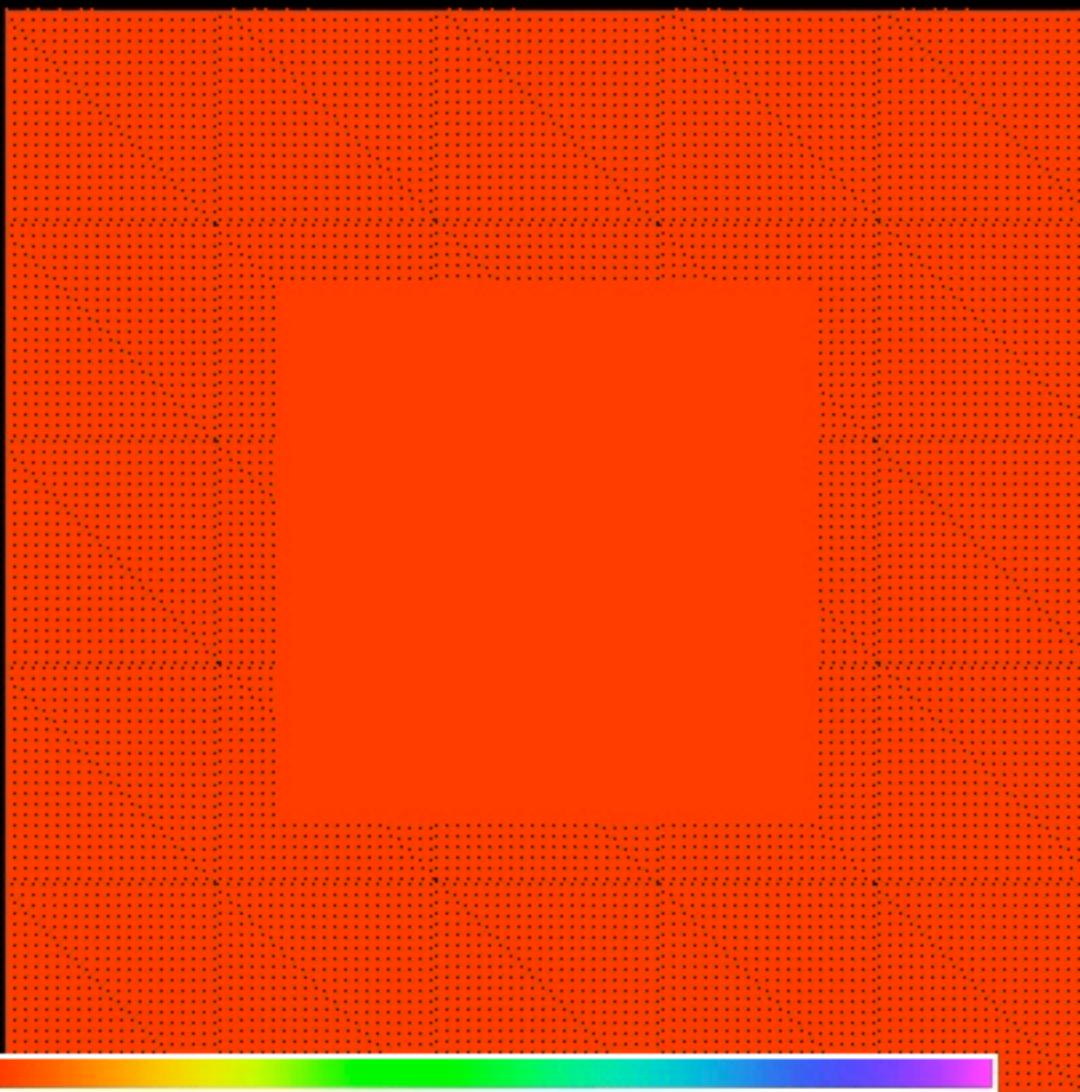
t=0.00





t=0.00

$\rho_0^{\text{no magnetic}} - \rho_0^{\text{magnetic}}$



0.00e+00

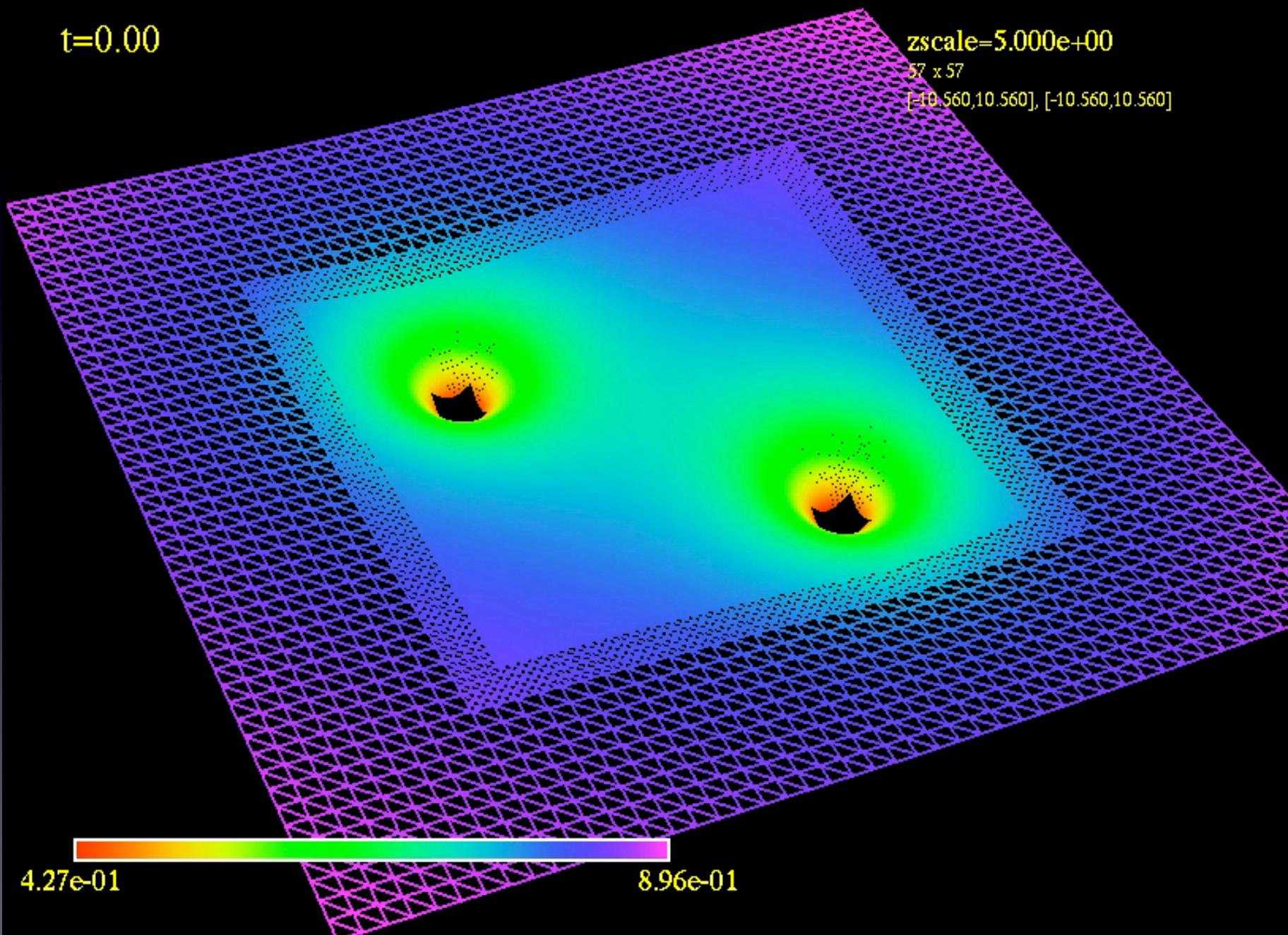
0.00e+00

t=0.00

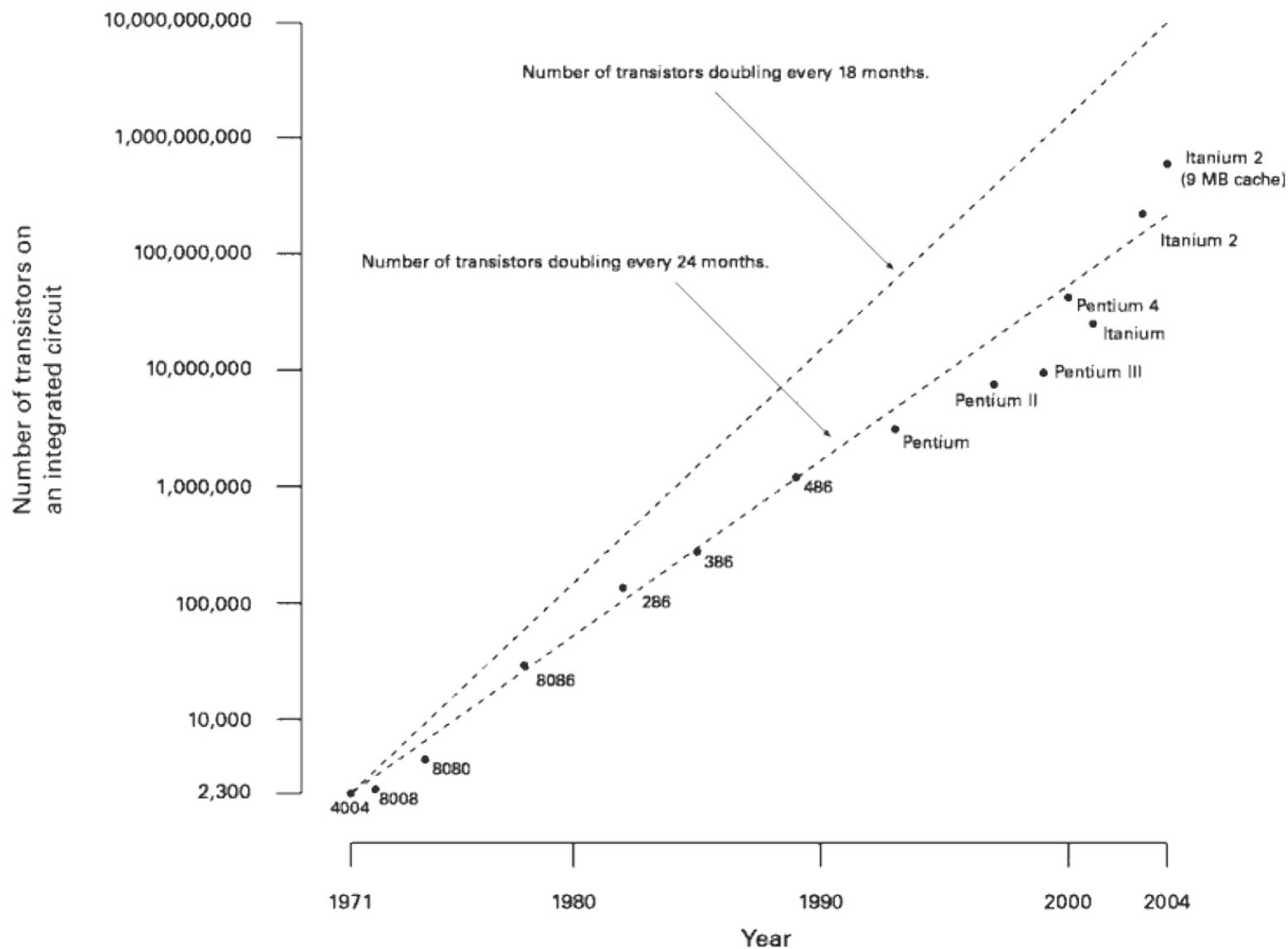
zscale=5.000e+00

57 x 57

[-10.560,10.560], [-10.560,10.560]



Moore's Law



Petascale in 2010(?)

Challenges with the current message passing model

- Heterogeneous computing components
 - Multicore, GPUs, Clearspeed SIMD, IBM Cell
- Latency
- Overhead
- Starvation
- Resource contention
- Fault tolerance

ParalleX Model for future computation

Thomas Sterling, CCT, LSU

- Locality (synchronous local resources)
- System wide global name space
- Multithreaded
- Parcels
- Local Control Objects
- Percolation
- Echo

Future directions

- Role of magnetic field in collapse
- General equations of state
- BH-NS
- Move towards petascale platforms
- Radiation hydrodynamics