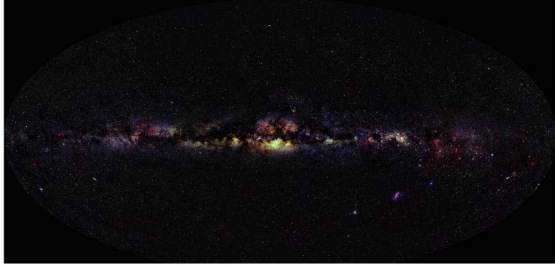


Chapter 19 Our Galaxy



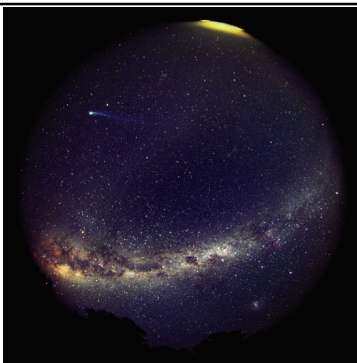
Agenda

- Announce:
 - Lunar Eclipse Saturday
 - Discuss GRB movie “Death Star”
- Ch 19– Our Galaxy
- Lab

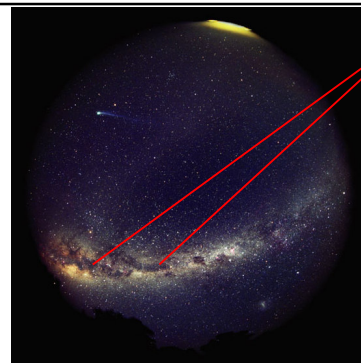
19.1 The Milky Way Revealed

- Our goals for learning
 - What does our galaxy look like?
 - How do stars orbit in our galaxy?

What does our galaxy look like?



The Milky Way galaxy appears in our sky as a faint band of light

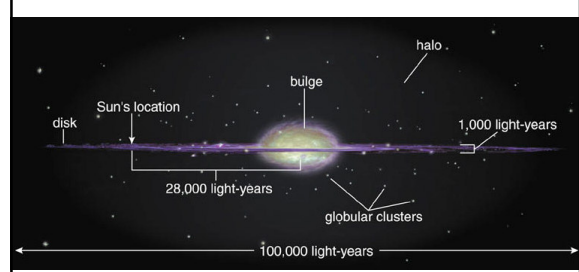


Dusty gas clouds obscure our view because they absorb visible light

This is the *interstellar medium* that makes new star systems

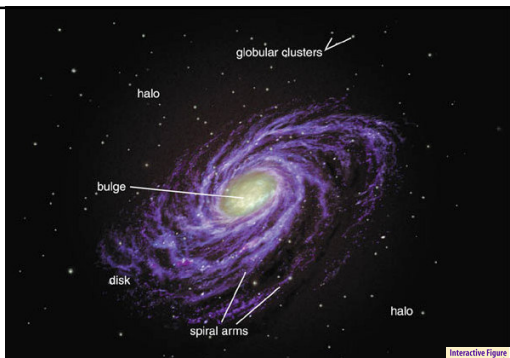


All-Sky View



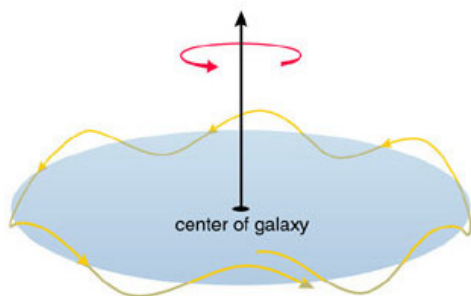
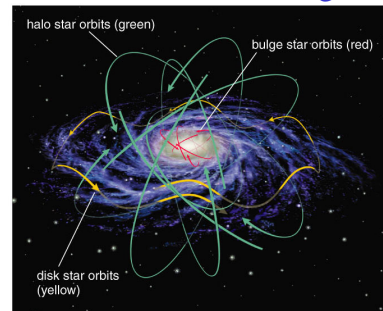
We see our galaxy edge-on

Primary features: disk, bulge, halo, globular clusters

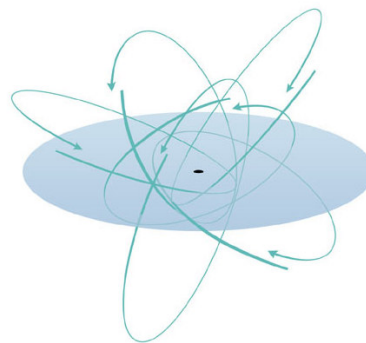


If we could view the Milky Way from above the disk, we would see its spiral arms

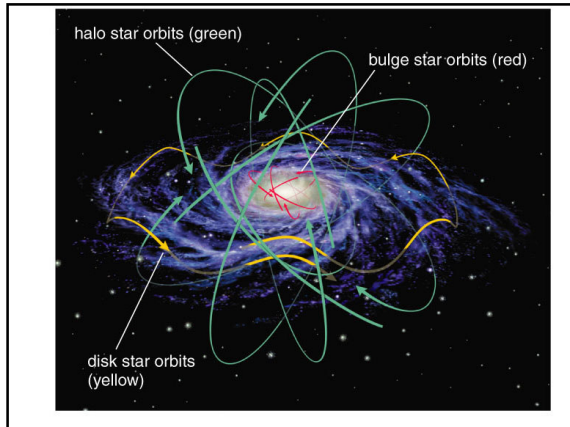
How do stars orbit in our galaxy?



Stars in the disk all orbit in the same direction with a little up-and-down motion



Orbits of stars in the bulge and halo have random orientations



Thought Question

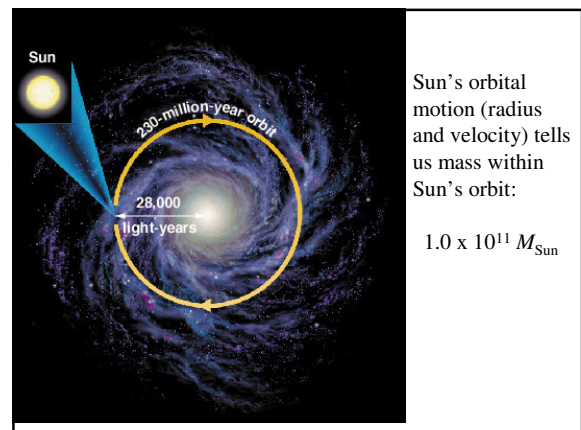
Why do orbits of bulge stars bob up and down?

- A. They're stuck to interstellar medium
- B. Gravity of disk stars pulls toward disk
- C. Halo stars knock them back into disk

Thought Question

Why do orbits of bulge stars bob up and down?

- A. They're stuck to interstellar medium
- B. Gravity of disk stars pulls toward disk**
- C. Halo stars knock them back into disk



Sun's orbital motion (radius and velocity) tells us mass within Sun's orbit:

$$1.0 \times 10^{11} M_{\text{Sun}}$$

Orbital Velocity Law

$$M_r = \frac{r \times v^2}{G}$$

- The orbital speed (v) and radius (r) of an object on a circular orbit around the galaxy tells us the mass (M_r) within that orbit

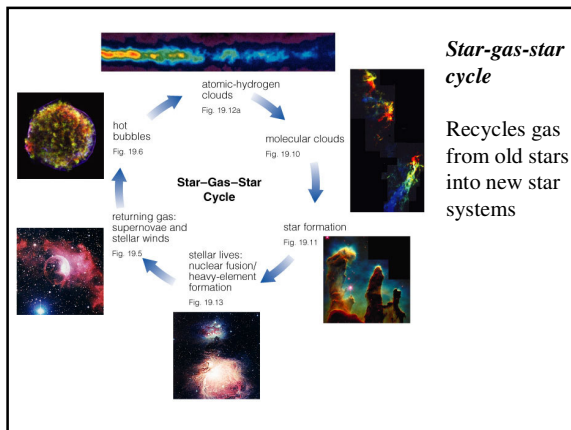
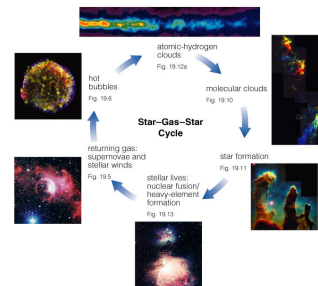
What have we learned?

- **What does our galaxy look like?**
 - Our galaxy consists of a disk of stars and gas, with a bulge of stars at the center of the disk, surrounded by a large spherical halo
- **How do stars orbit in our galaxy?**
 - Stars in the disk orbit in circles going in the same direction with a little up-and-down motion
 - Orbits of halo and bulge stars have random orientations

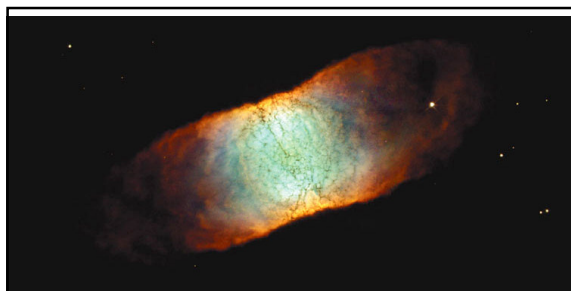
19.2 Galactic Recycling

- Our goals for learning
- How is gas recycled in our galaxy?
- Where do stars tend to form in our galaxy?

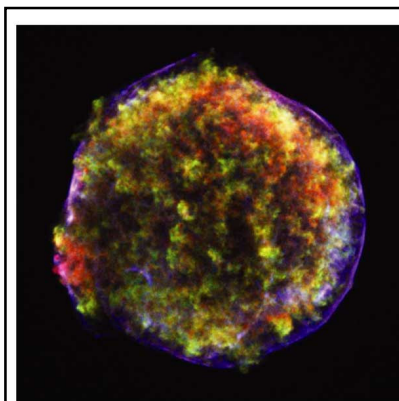
How is gas recycled in our galaxy?



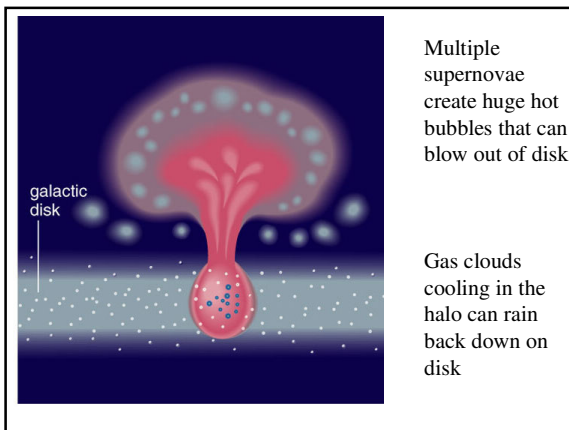
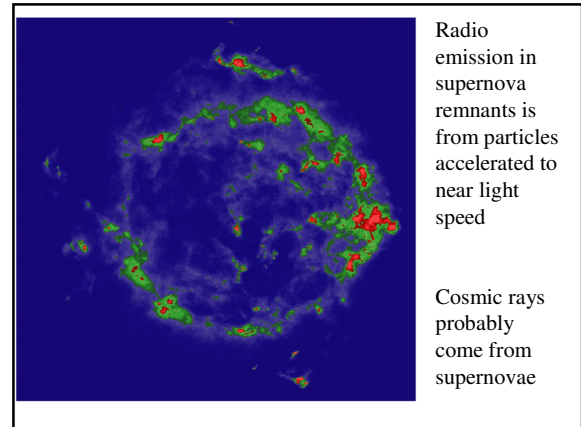
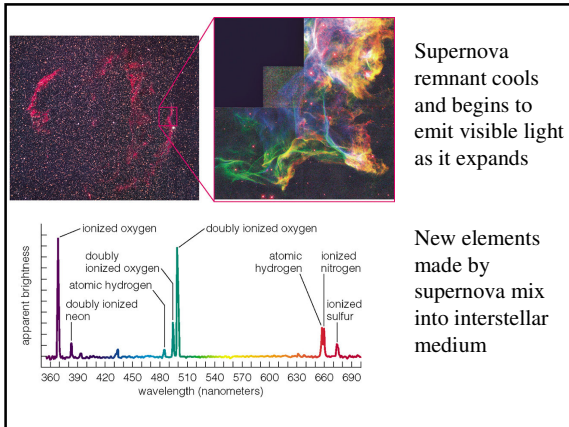
High-mass stars have strong stellar winds that blow bubbles of hot gas



Lower mass stars return gas to interstellar space through stellar winds and planetary nebulae

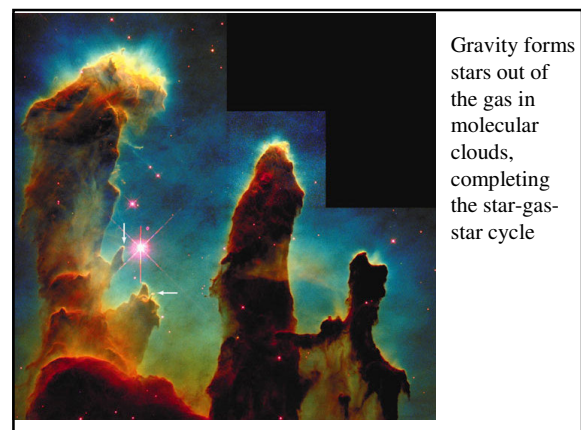
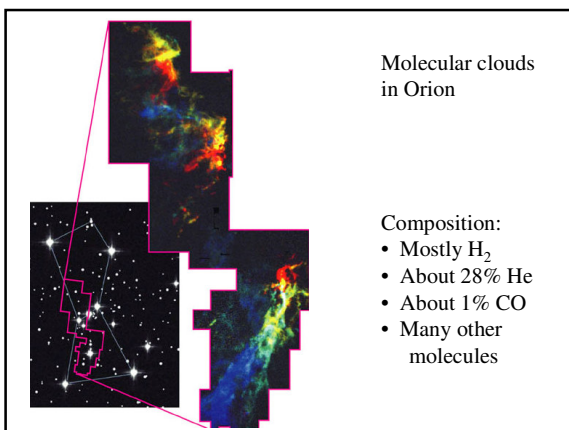


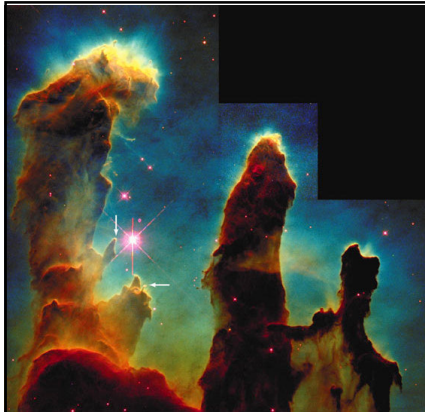
X-rays from hot gas in supernova remnants reveal newly-made heavy elements



Atomic hydrogen gas forms as hot gas cools, allowing electrons to join with protons

Molecular clouds form next, after gas cools enough to allow to atoms to combine into molecules





Radiation from newly formed stars is eroding these star-forming clouds

Summary of Galactic Recycling

- Gas Cools
- Stars make new elements by fusion
 - Dying stars expel gas and new elements, producing hot bubbles ($\sim 10^6$ K)
 - Hot gas cools, allowing atomic hydrogen clouds to form (~ 100 -10,000 K)
 - Further cooling permits molecules to form, making molecular clouds (~ 30 K)
 - Gravity forms new stars (and planets) in molecular clouds

Thought Question

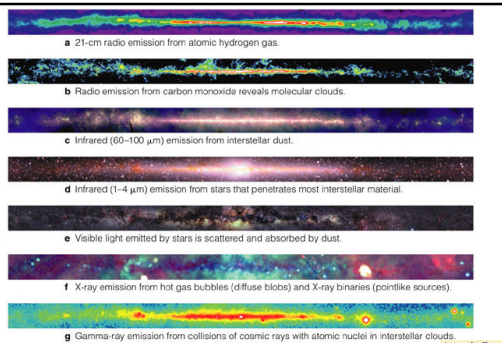
Where will the gas be in 1 trillion years?

- A. Blown out of galaxy
- B. Still recycling just like now
- C. Locked into white dwarfs and low-mass stars

Thought Question

Where will the gas be in 1 trillion years?

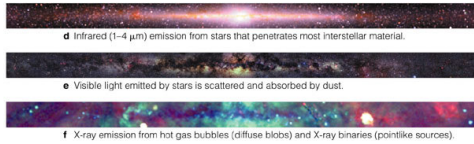
- A. Blown out of galaxy
- B. Still recycling just like now
- C. Locked into white dwarfs and low-mass stars**



We observe star-gas-star cycle operating in Milky Way's disk using many different wavelengths of light

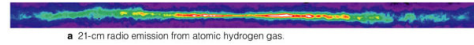


Infrared light reveals stars whose visible light is blocked by gas clouds

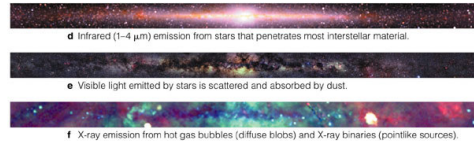


X-rays

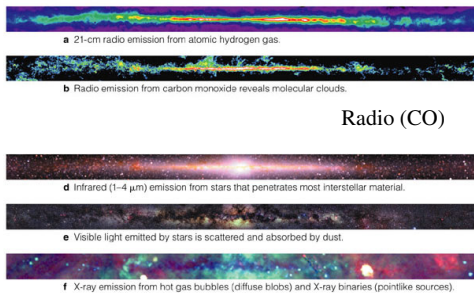
X-rays are observed from hot gas above and below the Milky Way's disk



Radio (21cm)

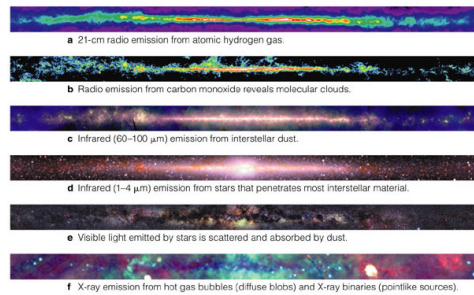


21-cm radio waves emitted by atomic hydrogen show where gas has cooled and settled into disk



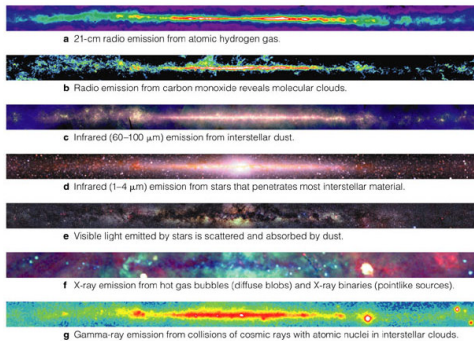
Radio (CO)

Radio waves from carbon monoxide (CO) show locations of molecular clouds



IR (dust)

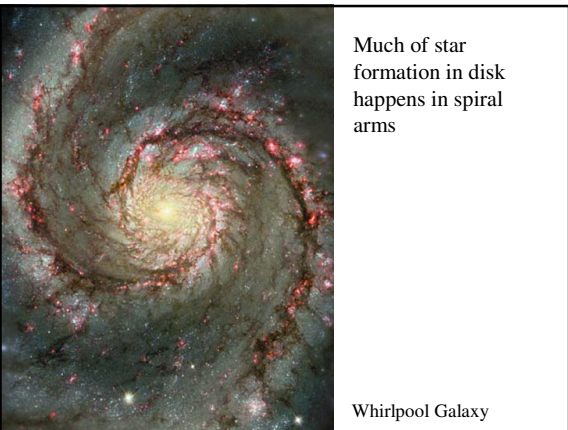
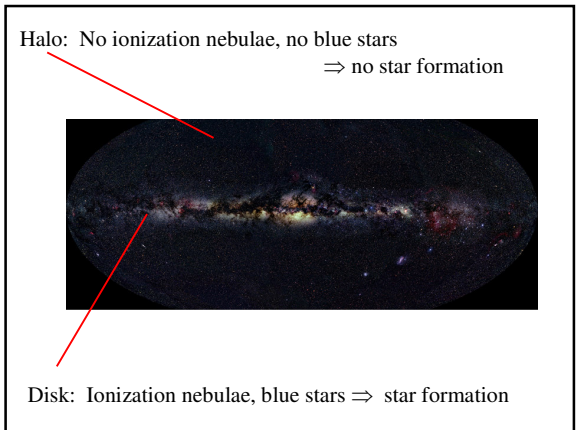
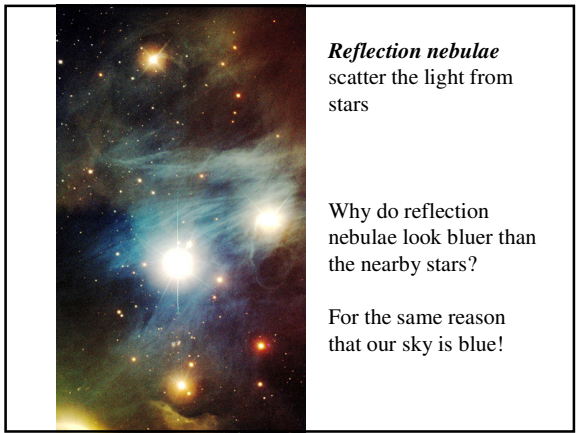
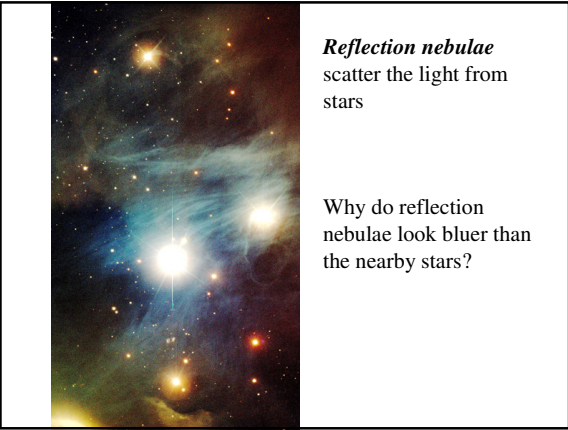
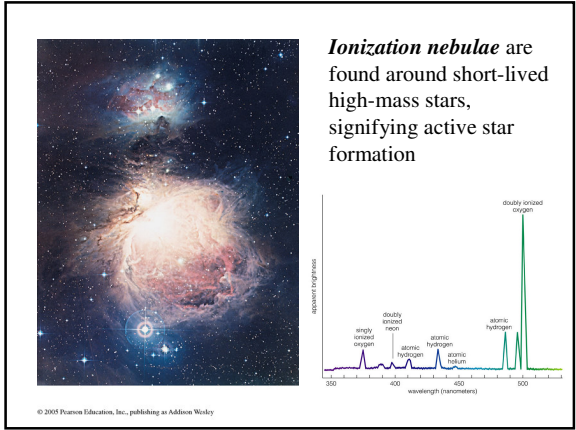
Long-wavelength infrared emission shows where young stars are heating dust grains

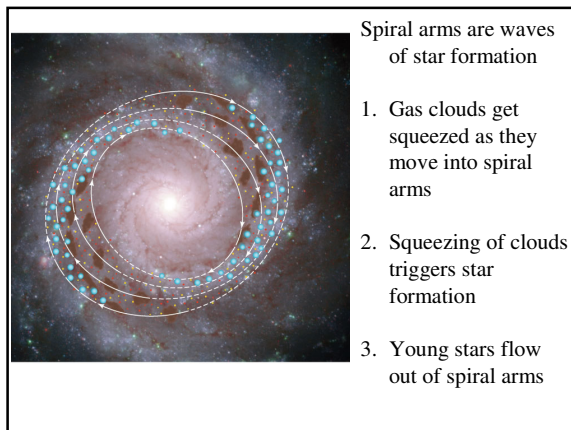
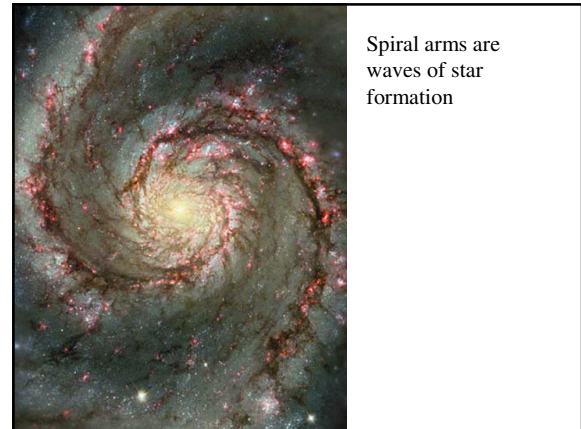
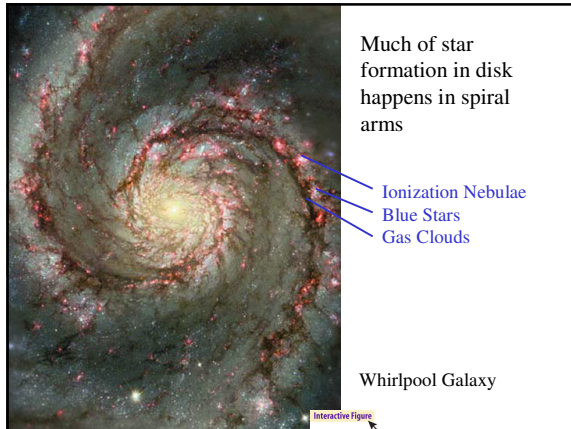


Gamma rays show where cosmic rays from supernovae collide with atomic nuclei in gas clouds

Where do stars tend to form in our galaxy?







What have we learned?


- **How is gas recycled in our galaxy?**
 - Gas from dying stars mixes new elements into the interstellar medium which slowly cools, making the molecular clouds where stars form
 - Those stars will eventually return much of their matter to interstellar space
- **Where do stars tend to form in our galaxy?**
 - Active star-forming regions contain molecular clouds, hot stars, and ionization nebulae
 - Much of the star formation in our galaxy happens in the spiral arms

19.3 The History of the Milky Way

- Our goals for learning
- What clues to our galaxy's history do halo stars hold?
- How did our galaxy form?


What clues to our galaxy's history do halo stars hold?

Halo Stars:
0.02-0.2% heavy elements (O, Fe, ...),
only old stars



Disk Stars:
2% heavy elements,
stars of all ages


Halo Stars:
0.02-0.2% heavy elements (O, Fe, ...),
only old stars



Disk Stars:
2% heavy elements,
stars of all ages

Halo stars
formed first,
then stopped

Halo Stars:
0.02-0.2% heavy elements (O, Fe, ...),
only old stars



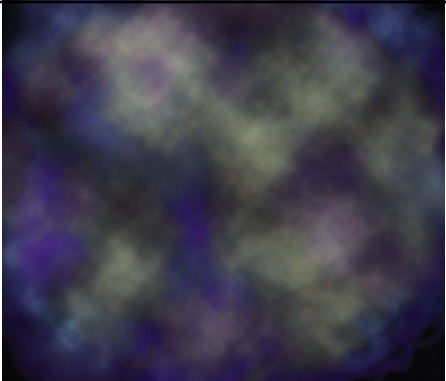
Disk Stars:
2% heavy elements,
stars of all ages

Halo stars
formed first,
then stopped

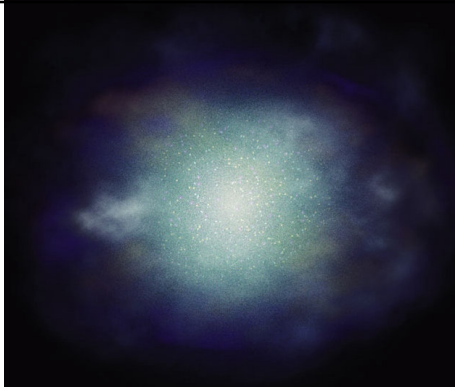
Disk stars
formed later,
kept forming

How did our galaxy form?

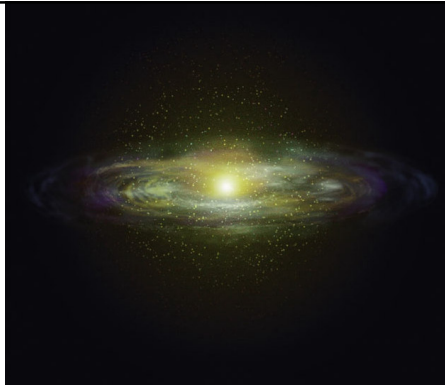




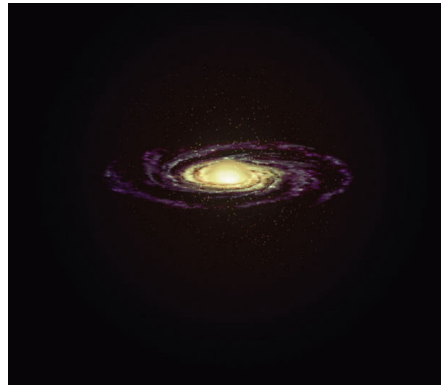
Our galaxy probably formed from a giant gas cloud



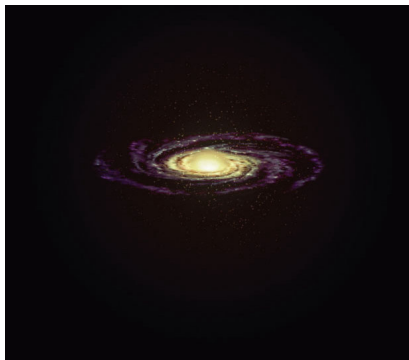
Halo stars formed first as gravity caused cloud to contract



Remaining gas settled into spinning disk

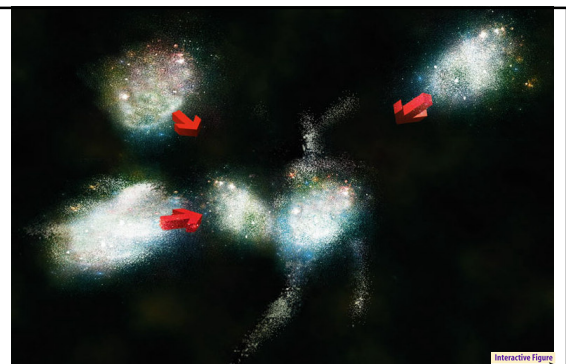


Stars continuously form in disk as galaxy grows older



Stars continuously form in disk as galaxy grows older

Warning: This model is oversimplified



Detailed studies: Halo stars formed in clumps that later merged

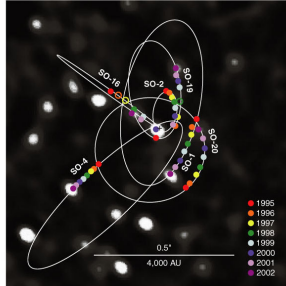
What have we learned?

- What clues to our galaxy's history do halo stars hold?
 - Halo stars are all old, with a smaller proportion of heavy elements than disk stars, indicating that the halo formed first
- How did our galaxy form?
 - Our galaxy formed from a huge cloud of gas, with the halo stars forming first and the disk stars forming later, after the gas settled into a spinning disk

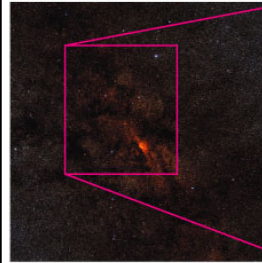
19.4 The Mysterious Galactic Center

- Our goals for learning
- What lies in the center of our galaxy?

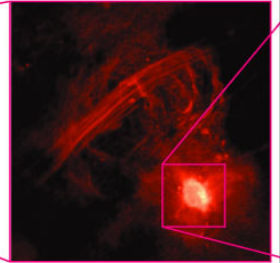
What lies in the center of our galaxy?



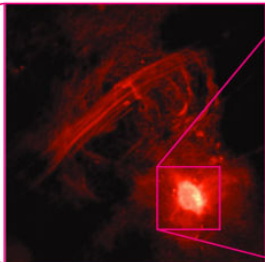
Infrared light from center



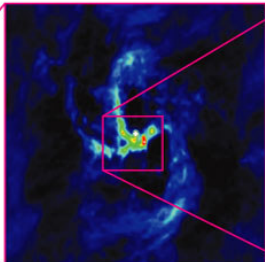
Radio emission from center



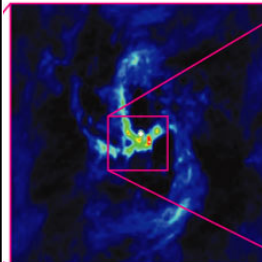
Radio emission from center



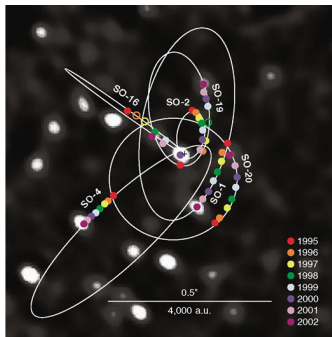
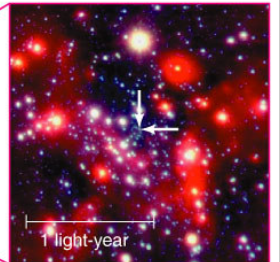
Swirling gas near center



Swirling gas near center

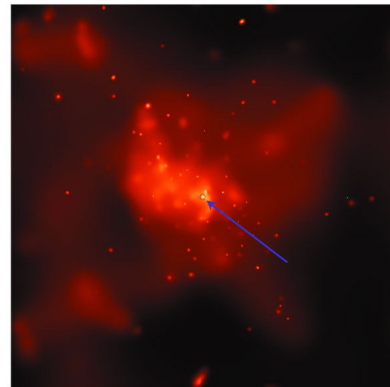


Orbiting star near center



Stars appear to be orbiting something massive but invisible ... *a black hole?*

Orbits of stars indicate a mass of about 4 million M_{Sun}



X-ray flares from galactic center suggest that tidal forces of suspected black hole occasionally tear apart chunks of matter about to fall in

What have we learned?

- What lies in the center of our galaxy?
 - Orbits of stars near the center of our galaxy indicate that it contains a black hole with 4 million times the mass of the Sun